

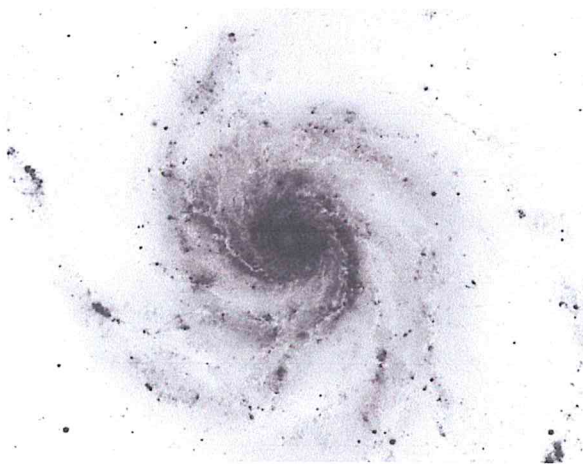
Podane so svetlobne krivulje v spektralnih območjih B, V in R za dve supernovi tipa Ia, ki so ju astronomi opazovali v dveh različnih galaksijah. Na abscisi grafov je čas v mesecih/dnevih, na ordinati pa navidezne magnitude v pripadajočih spektralnih območjih. Na fotografijah (v negativu) sta galaksiji, v katerih sta zasvetili supernovi. V preglednici so njune ekvatorialne koordinate.

Izračunaj oddaljenost obeh galaksij, če veš, da je absolutna magnituda supernov Ia v območju V, ko je njihov sij največji, -19.

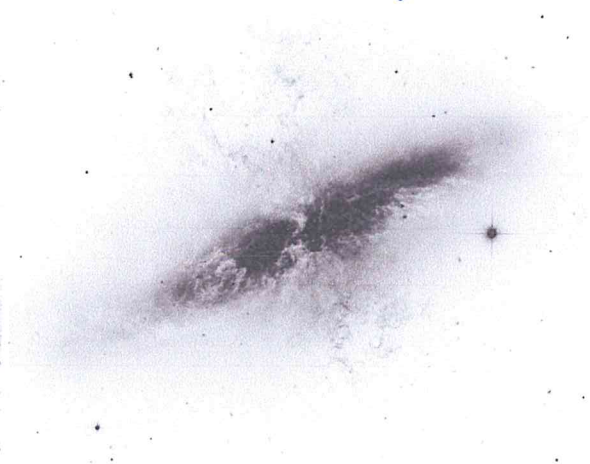
STA ISTI  
SPEKTRALNI  
KLAS

galaksija	$\alpha$	$\delta$
1	14 <sup>h</sup> 03 <sup>m</sup>	+54°21'
2	09 <sup>h</sup> 56 <sup>m</sup>	+69°41'

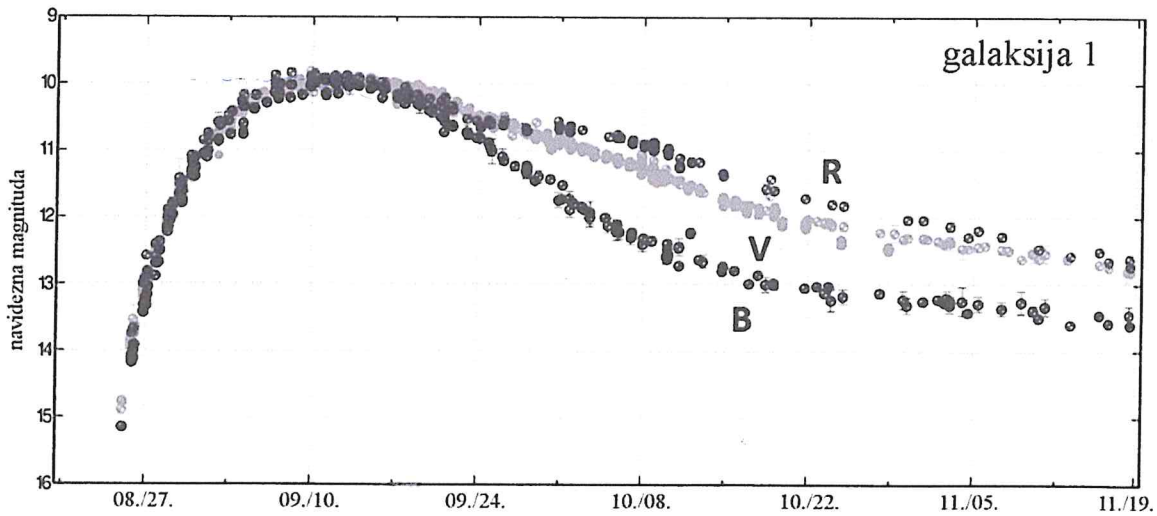
veja ekstinkcija

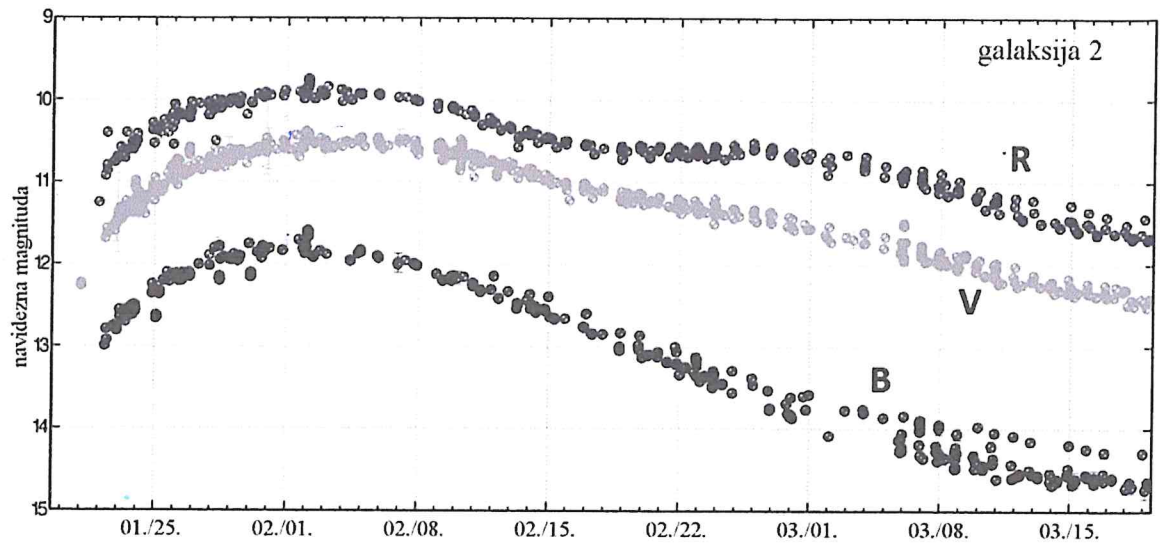


galaksija 1



galaksija 2





GALAXY 1:

$V_{max} = +10^m$

$\frac{F_M}{F} = \frac{r^2}{R_M^2} = 10^{-0.14(M-m)}$

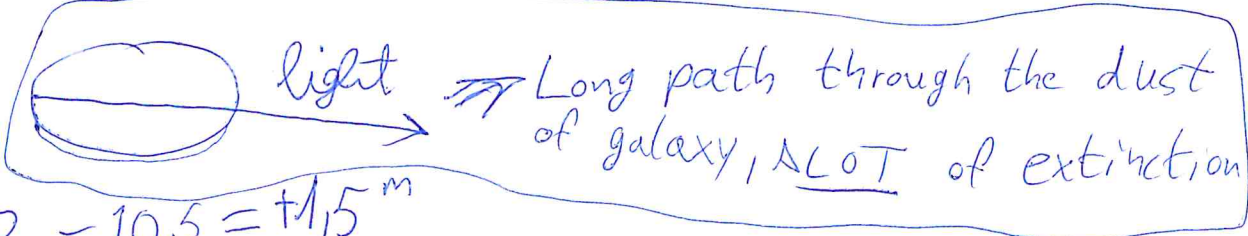
$\Rightarrow r = 10^{0.12(m-M)}$

short path through the dust of galaxy, LITTLE EXTINCTION  $\rightsquigarrow$  since galaxies are very thin discs, we can take it to be  $\approx 0^m$ .

$B_1 - V_1 = M_B - M_V = 0^m$  !!!

$R_M = 10^{0.12(10 - (-19))} = 10^{2.28} = 10^{6.18} \text{ pc}$

GALAXY 2:



$B_2 - V_2 = 12 - 10.5 = +1.5^m$

$E_{B-V} = (M_B - M_V) + (B - V) = B - V = +1.5^m$

0 (shown in Galaxy 1 graph)  $\rightarrow$  because  $B_1 - V_1 = M_B - M_V$ , since in Galaxy 1 there is no extinction

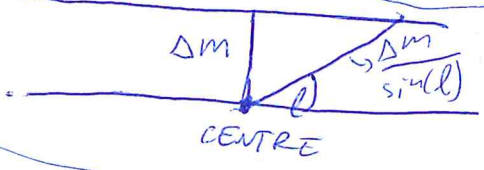
$\frac{E_V}{E_{B-V}} \approx 3 \Rightarrow E_V = 3 E_{B-V} = +4.5^m$

$V_{ACTUAL_2} = V_{MAX_2} - 4.5^m = 10.5^m - 4.5^m = 6^m$

$r_2 = 10^{0.12(V_{ACTUAL_2} - M_V)} = 10^{0.12(6 - (-19))} = 10^{2.28} = 10^6 \text{ pc}$

If we knew galactic latitude  $l$  of both galaxies, we could also estimate extinction due to OUR GALAXY:

OUR GALAXY:



$\Rightarrow$  extinction in our galaxy  $\propto \frac{1}{\sin(l)}$  if  $l \rightarrow$  galactic latitude

We could determine  $l$  of galaxy 1 and 2 from their  $\alpha$  and  $\delta$ ; if we knew equatorial coordinates of galactic centre.

We could estimate additional extinction due to intergalactic dust, if we knew, how many magnitudes per parsec extinction it causes:

$\Delta E_{INTER} = k \cdot r$ , where  $k$  is extinction constant.